

Gravitationally induced decoherence: from theoretical models to signatures in neutrino oscillations

Max Joseph Fahn

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Theory Seminar - Bologna, Jan 29, 2025

ALMA MATER STUDIORUM ~ UNIVERSITÀ DI BOLOGNA



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Quantum Theory (QT) General Relativity (GR)

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Isolated Quantum Systems

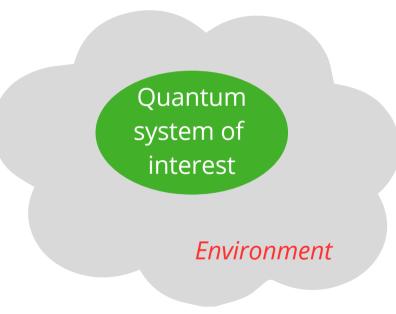


Liouville-von Neumann equation

$$rac{\partial
ho_S(t)}{\partial t} = -i[H_S(t),
ho_S(t)]$$



Open Quantum Systems



Master equation

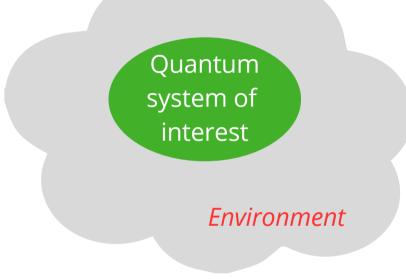
$$rac{\partial
ho_S(t)}{\partial t} = -i[H_S(t) {+} rac{H_{add}(t)}{H_{add}(t)},
ho_S(t)] {+} \mathcal{D}[
ho_S(t)]$$

New processes encoded in red terms:

- Energy shifts / Renormalisation
- Dissipation
- Decoherence
 - \checkmark Vanishing of off-diagonal elements of $ho_S(t)$
 - \rightarrow no interference
 - \rightarrow "classicalisation"



Open Quantum Systems



Master equation

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ho_S(t)}{\partial t} = -i[H_S(t) {+} rac{H_{add}(t)}{H_{add}(t)},
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New processes encoded in red terms:

- Energy shifts / Renormalisation
- Dissipation
- Decoherence
- Without gravity: Frequently used (e.g. Quantum Optics, QED, Solid State Physics, ...)
- With (quantised) gravity as environment: Rather new
 - \leftrightarrow Need some Quantum Gravity Theory as basis



Overview of the talk

Part I: Fieldtheoretical master equation

Part II: Oneparticle projection of the master equation "A gravitationally induced decoherence model using Ashtekar variables" (2022); *MJF, K. Giesel, M. Kobler* (CQG)

"Gravitationally induced decoherence of a scalar field: investigating the one-particle sector and its interplay with renormalisation" (2024); *MJF, K. Giesel* (arXiv)

Part III: Application to neutrino oscillations "Understanding gravitationally induced decoherence parameters in neutrino oscillations using a microscopic quantum mechanical model" (2024); A. Domi, T. Eberl, MJF, K. Giesel, L. Hennig, U. Katz, R. Kemper, M. Kobler (JCAP)



Part I

Field-theoretical master equation



"A gravitationally induced decoherence model using Ashtekar variables" (2022); *MJF*, *K. Giesel, M. Kobler* (CQG)



Master equations with gravity

[Anastopoulos, Hu 2013], [Blencowe 2013], [Oniga, Wang 2016], [Lagouvardos, Anastopoulos 2021], [Fahn, Giesel, Kobler 2022]

Classical system

$$S=S_{EH}[g_{ij},P^{ij}]+S_{\Phi_i}[\Phi_i,\Pi^j,g_{ij},P^{ij}]$$
 Coupling of gravity and matter by GR



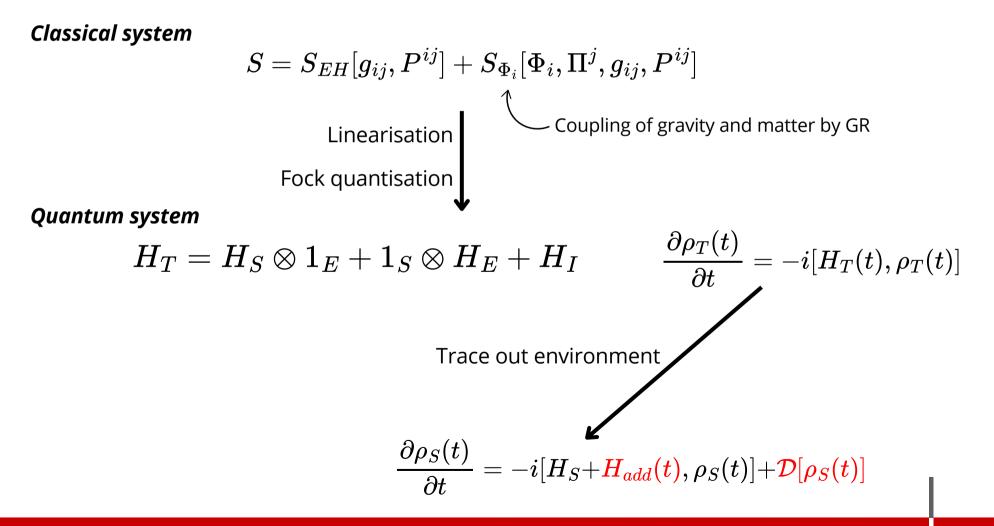
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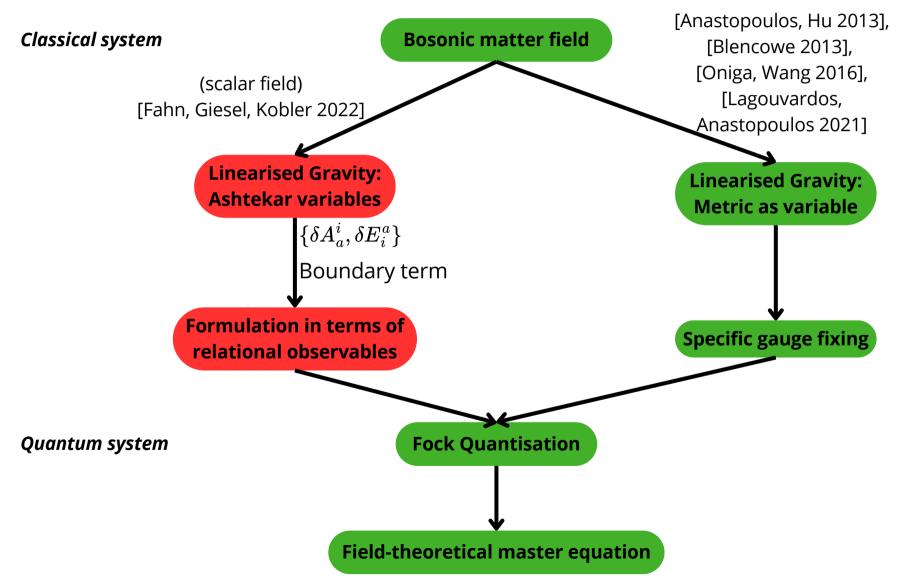
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Procedure Part I



Max Joseph Fahn, Theory Seminar Bologna



[Rovelli 1991], [Rovelli 1991], [Vytheeswaran 1994], [Rovelli 2002], [Dittrich 2006], [Dittrich, 2007]

- Method to deal with gauge freedom in physical systems and to formulate dynamics
- Basic idea: Describe evolution in terms of other objects; here: geometric clocks



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- Construct (weakly) gauge invariant observables up to second order in κ using the observable map

$${\mathcal O}_{f,\{T\}}(au^I) = \left[\exp(\xi^I\{C_I,\cdot\}) \cdot f
ight] \Big|_{\xi^I:=T^I- au^I}$$

with first class constraints $\{C_I = \delta C_I + \delta^2 C_I\}$, clocks $\{T^I = \delta T^I\}$

• Construct geometrical clocks that fulfill $\{\delta T^I, \delta C_J\} = \delta^I_J$



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- Construct geometrical clocks that fulfill $\{\delta T^I, \delta C_J\} = \delta^I_J$
- Use dual map to make reference fields commute mutually, $\{\delta T^{I}, \delta T^{J}\} = 0$
- Canonical Transformation:

$$(\phi,\pi), \quad (\delta A_a^i, \delta E_i^a) \longrightarrow (\delta T^I, \delta C_J), \quad \underbrace{(\mathcal{O}_{\phi}, \mathcal{O}_{\pi}), (\mathcal{O}_{\delta \mathcal{A}_a^i}, \mathcal{O}_{\delta \mathcal{E}_i^a})}_{\text{Closed subalgebra}}$$



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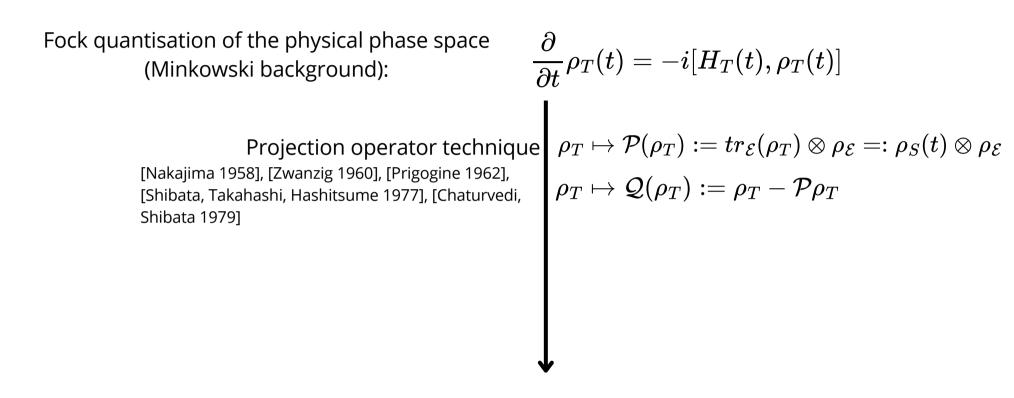
Separation of the phase space into physical and unphysical degrees of freedom



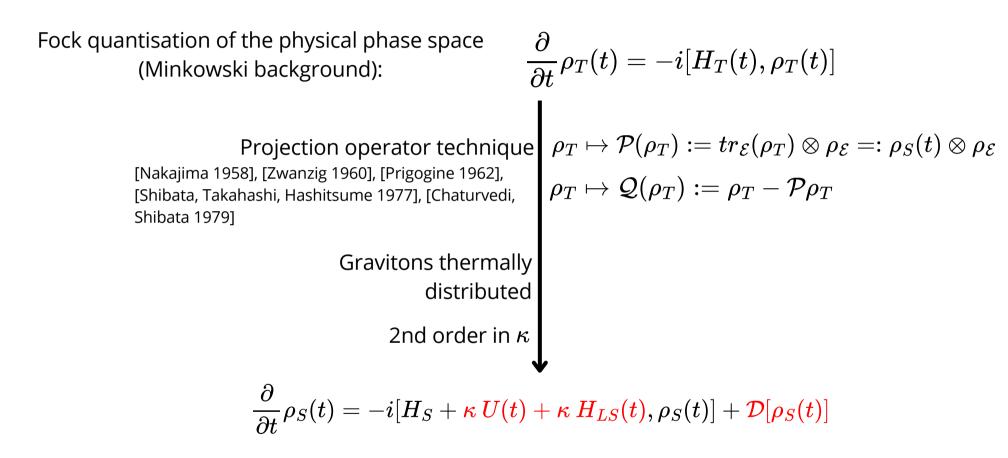
Fock quantisation of the physical phase space (Minkowski background):

$$rac{\partial}{\partial t}
ho_T(t)=-i[H_T(t),
ho_T(t)]$$

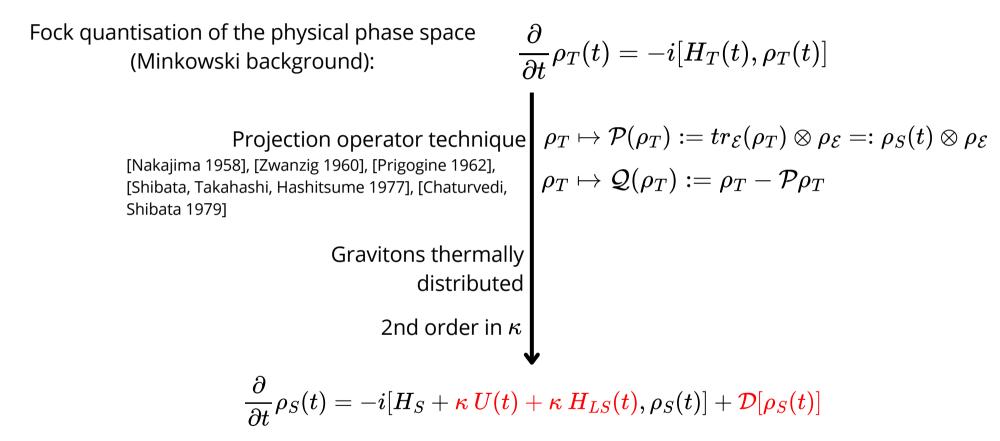












Challenges:

- Structure still very complicated \rightarrow further approximations required for solution
- Connect to experiments \rightarrow dynamics of a single scalar particle
- Extract physics \rightarrow renormalisation



Part II

One-particle projection of the master equation

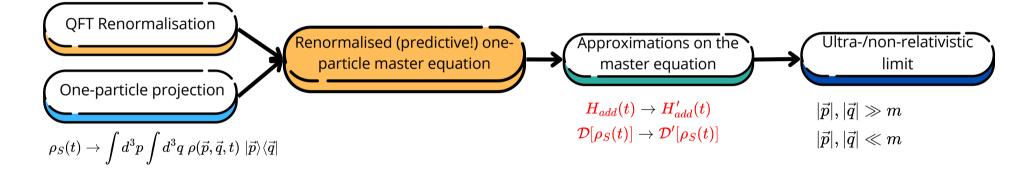


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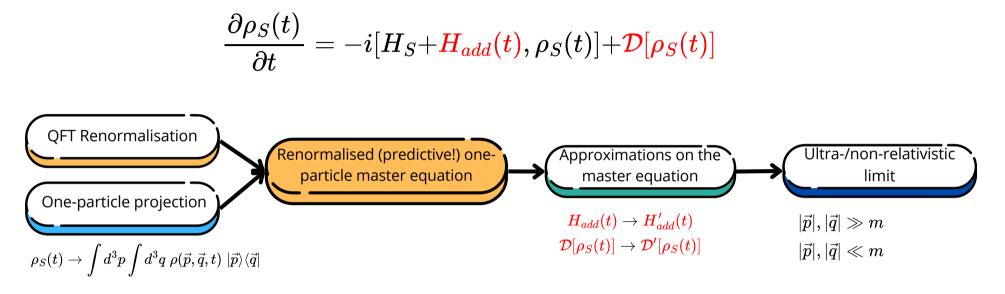
Plan for Part II

$$rac{\partial
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ho_S(t)]$$





Plan for Part II



With QM/ Without Renormalisation: [Anastopoulos, Hu 2013], [Oniga, Wang 2016], [Lagouvardos, Anastopoulos 2021]



Interpretation of divergences? Physical effect of the approximations?



QFT Renormalisation

- Some terms of the master equation exhibit logarithmic UV divergences
- Individual terms of the one-particle master equation can be connected to Feynman diagrams [Burrage et al. 2019]



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- Feynman rules from the initial Hamiltonian (Relational Observables):

• UV divergence comes from vacuum self-energy: $\underbrace{u \quad \int_{u-k}^{k} u}_{u-k}$



QFT Renormalisation

- Some terms of the master equation exhibit logarithmic UV divergences
- Individual terms of the one-particle master equation can be connected to Feynman diagrams [Burrage et al. 2019]
- Feynman rules from the initial Hamiltonian (Relational Observables):

- UV divergence comes from vacuum self-energy: $\underbrace{u \quad \sqrt{k} \\ u k}{k}$
- Restore covariance (similar to QED in Coulomb gauge) [Tong QFT script], [Weinberg QFT book]
- Renormalise self-energy loop (dimensional regularisation and small artificial graviton mass to cure IR divergence; on-shell renormalisation scheme)
- Effect: Renormalisation removes all vacuum terms from the dissipator



(Further) Approximations

Typical aim: obtain a completely positive master equation

 \rightarrow further approximations:

Markov approximation

Idea: Environment forgets rapidly the influence of the system

Justified for ultra-relativistic particles

Rotating wave approximation

Idea: "Coarse graining": Detectors cannot measure arbitrarily fast oscillations



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Dissipator has Lindblad form, no Lamb-shift left*:

$$egin{aligned} &rac{\partial}{\partial t}
ho_1(t)=-i[H_S,
ho_1(t)]+\mathcal{D}[
ho_1(t)]\ &\mathcal{D}[
ho_1]=\kappa\sum_{r\in\{+,-\}}\int_{\mathbb{R}^3}rac{d^3k}{(2\pi)^2}\delta(k)rac{N(k)}{\Omega_k}\Big(L_r(ec{k})
ho_1L_r(ec{k})^\dagger-rac{1}{2}\Big\{
ho_1,L_r(ec{k})^\dagger L_r(ec{k})\Big\}\Big) \end{aligned}$$



Specific limits

Non-relativistic limit:

$$rac{\partial}{\partial t}
ho(ec p,ec q,t)=-i
ho(ec p,ec q,t)\left(rac{ec p^2}{2m}-rac{ec q^2}{2m}
ight)-rac{2\kappa k_BT}{3m^2}igg[rac{1}{3}(ec p^4+ec q^4)-rac{1}{3}ec p^2ec q^2\left(1+\cos^2(\gamma)
ight)igg]
ho(ec p,ec q,t)$$

Agrees approx. with [Anastopoulos, Hu 2013]

Ultra-relativistic limit:

$$rac{\partial}{\partial t}
ho(ec{p},ec{q},t) = -i
ho(ec{p},ec{q},t)\;(|ec{p}| - |ec{q}|) - rac{\kappa k_BT}{4\pi} igg[rac{1}{3}(ec{p}^2 + ec{q}^2) + |ec{p}||ec{q}| \left(rac{1}{3} - \cos^2(\gamma)
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As operator equation in 1D:

$$\frac{\partial}{\partial t}\hat{\rho}(t) = -i[\hat{H}_{S},\hat{\rho}(t)] + \frac{\kappa k_{B}T}{6\pi} \left(\hat{H}_{S}\hat{\rho}(t)\hat{H}_{S} - \frac{1}{2}\left\{\hat{H}_{S}^{2},\hat{\rho}(t)\right\}\right) \quad \text{[Assumed approximately } \vec{p} \parallel \vec{q} \text{]}$$
Solution in energy basis assuming discrete energy levels:
$$(t) = -i(E_{i}-E_{i})t - \frac{\kappa}{2\pi^{2}}(E_{i}-E_{i})^{2}t$$

$$ho_{ij}(t)=
ho_{ij}(0)e^{-i(E_i-E_j)t-rac{\kappa}{12\pieta}(E_i-E_j)^2t}$$



Part III

Application to neutrino oscillations



"Understanding gravitationally induced decoherence parameters in neutrino oscillations using a microscopic quantum mechanical model" (2024); A. Domi, T. Eberl, MJF, K. Giesel, L. Hennig, U. Katz, R. Kemper, M. Kobler (JCAP)



Model of interest

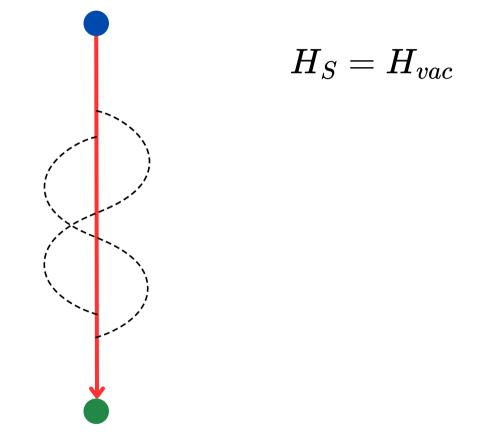


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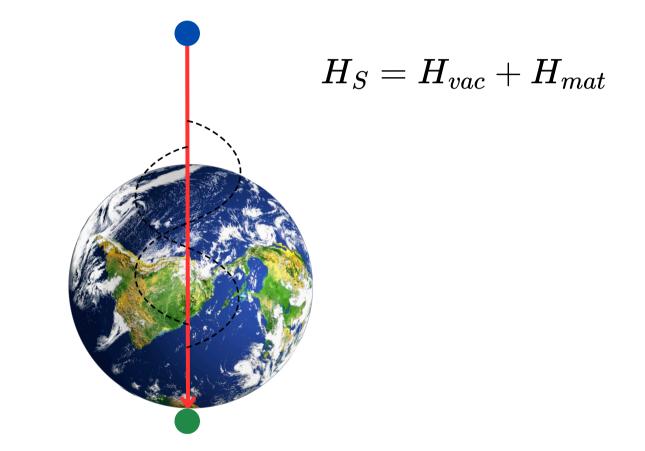
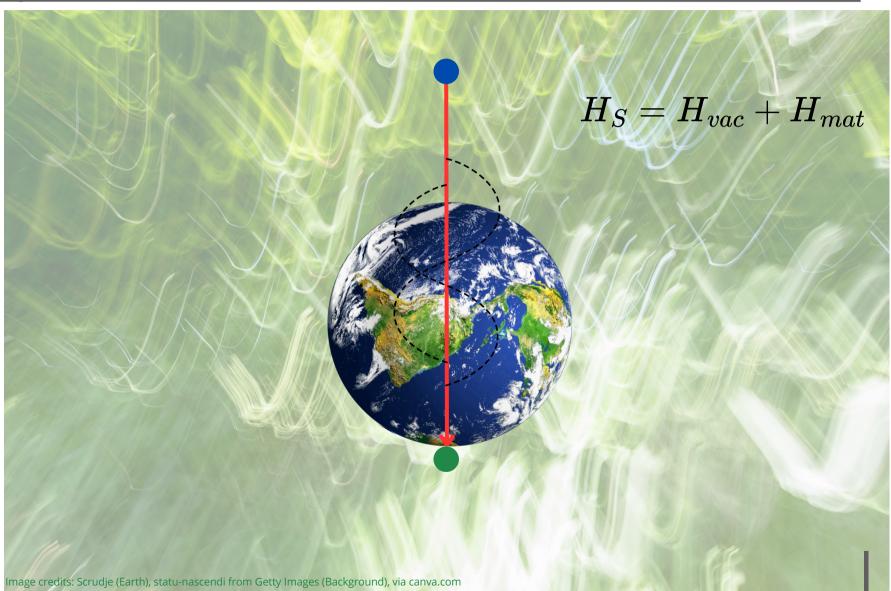


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Model of interest





Phenomenological models

Phenomenological quantum mechanical models

[Lisi, Marrone, Montanino, 2000], [Benatti, Floreanini 2000], [Gago, Santos, Teves , Zukanovich Funchal 2002], [Morgan, Winstanley, Brunner, Thompson 2006], [Guzzo, Holanda, Oliveira 2016], [Coelho, Mann 2017], [Coloma, Lopez-Pavon, Martinez-Soler, Nunokawa 2018], [Gomes, Gomes, Peres 2023], ...

Typical procedure:

- Lindblad equation
- Solution (in effective mass basis):

$$\widetilde{
ho}_{ij}(t) = \widetilde{
ho}_{ij}(0) \cdot e^{-rac{i}{\hbar} \left(\widetilde{H}_i - \widetilde{H}_j
ight) t - \Gamma_{ij} t} \qquad \Gamma_{ij} = \gamma_{ij} \cdot \left(rac{E}{E_0}
ight)^n$$

• Gravitationally induced decoherence manifests as damping of the neutrino oscillations



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• Gravitationally induced decoherence manifests as damping of the neutrino oscillations

Based on Part I and Part II, can we come up with a microscopic model to resolve these unknown parameters?



Quantum mechanical toy model

Hamiltonian, inspired by [Xu, Blencowe 2022]:

$$\hat{H} = ig(\underbrace{\hat{H}_S^{(0)} + \hat{H}_S^{(C)}}_{\hat{H}_S} ig) \otimes 1_\mathcal{E} + 1_S \otimes ig[\underbrace{rac{1}{2} \sum_{i=1}^N ig(\hat{p}_i^2 + \omega_i^2 \hat{q}_i^2 ig)}_{\hat{H}_\mathcal{E}} ig] - \underbrace{\hat{H}_S \otimes \sum_{i=1}^N g_i \hat{q}_i}_{\hat{H}_{ ext{int}}}$$



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Part II
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Major steps:

- Small coupling constant, environment follows Gibbs state
- Continuum limit of the environment
- Markov approximation
- Renormalisation



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Solution of the master equation (in effective mass basis):

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ight)t-rac{4\eta^2}{\hbar^3eta(T)}\left(\widetilde{H}_i-\widetilde{H}_j
ight)^2t}$$

(same form as one-particle master equation from Part II in ultra-relativistic limit)



Comparison to the phenomenological models

Solution of the master equation:

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Solution for phenomenological models:

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Results:

- Expression for phenomenological parameters
- Energy dependence fixed
- Match to phenomenological models only in vacuum possible



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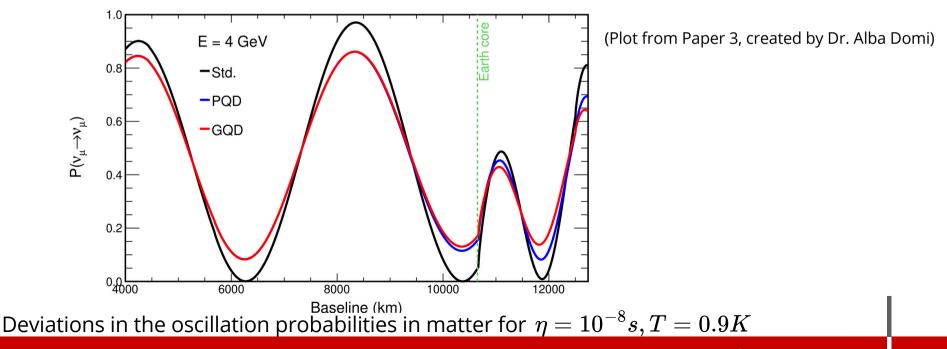
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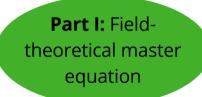
(Quantised linearised) Gravity as environment in Open Quantum Systems Main model: Scalar field coupled to linearised gravity in Ashtekar variables

Part I: Field-
theoretical master
equation

- Relational formalism:
 - Observables
 - Notion of dynamics
 - Physical time and length
- Projection operator technique



(Quantised linearised) Gravity as environment in Open Quantum Systems Main model: Scalar field coupled to linearised gravity in Ashtekar variables

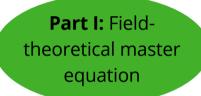


Part II: Oneparticle projection of the master equation

- Relational formalism:
 - Observables
 - Notion of dynamics
 - Physical time and length
- Projection operator technique
- One-particle projection
- Renormalisation
- Markov and Rotating wave approximations and their physical effects
- Application to non- and ultra-relativistic scalar particles



(Quantised linearised) Gravity as environment in Open Quantum Systems Main model: Scalar field coupled to linearised gravity in Ashtekar variables



Part II: Oneparticle projection of the master equation

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Part III: Application to neutrino oscillations

- QM toy model:
 - Renormalisation and Markov approximation
 - Resolution of the decoherence parameters in the phenomenological models; match only in vacuum



Outlook I

- Similar analysis for different matter systems (matter gauge degrees, different quantisation and renormalisation, ...) [Fahn, Giesel, Kemper, soon]
- Different gravitational constituent/background (cosmological, black hole, ...)
- Different quantisation of gravity [Giesel, Kobler 2022]



Outlook I

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- Different gravitational constituent/background (cosmological, black hole, ...)
- Different quantisation of gravity [Giesel, Kobler 2022]
- Analysis of the validity of Markov and Rotating Wave approximation
- Toy model \rightarrow favourable energy ranges and dependencies for phenomenological models
- Modification of the phenomenological models in matter
- Sensitivity analyses

Thank you for your attention!

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